

Observations and Analysis of the Mars 2001 Global Dust Storm

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Overview

- Goal: To better understand the dynamical processes involved in the origin, evolution, and decay of the storm
- Datasets include: Mars Orbital Camera (MOC) daily global maps, Thermal Emission Spectrometer (TES), Mars Horizon Sensor Assembly (MHSA).

Notes: Only a few representative plots are included in this presentation – please see complete animations at:
<http://humbabe.arc.nasa.gov/~noble/>

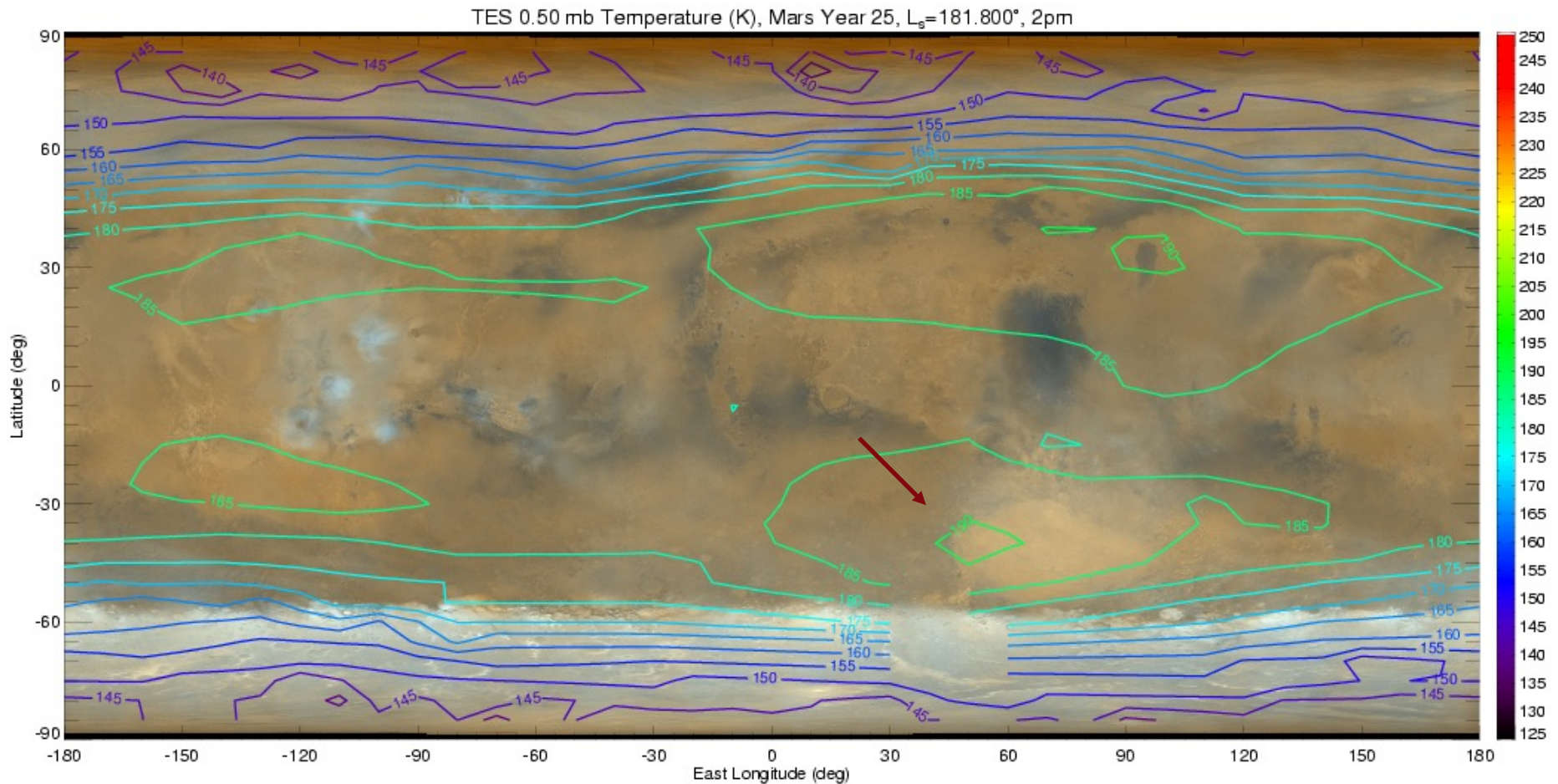
Comparison of datasets

- MOC, TES, MHSA maps are not true synoptic maps but are mosaics of 12 sun-synchronous swaths stitched together.
- MHSA and TES temperature data are quite consistent with each other, especially regarding the initiation and propagation of the wave one feature which reaches peak amplitude at $\sim L_s=187^\circ$
- TES data has a stronger horizontal temperature gradient than the MHSA data. This is due to the deep weighting function of the MHSA data (averaged over $\sim 10-40$ km).

Pre storm activity: $L_s=170^\circ-L_s=182^\circ$

- At least seven pulses, which came every 2 – 3 sols, led up to initiation of the global dust storm which began in Hellas. (Cantor, 2004)
- These pulses may be the signature of traveling baroclinic eddies since this is their typical period.

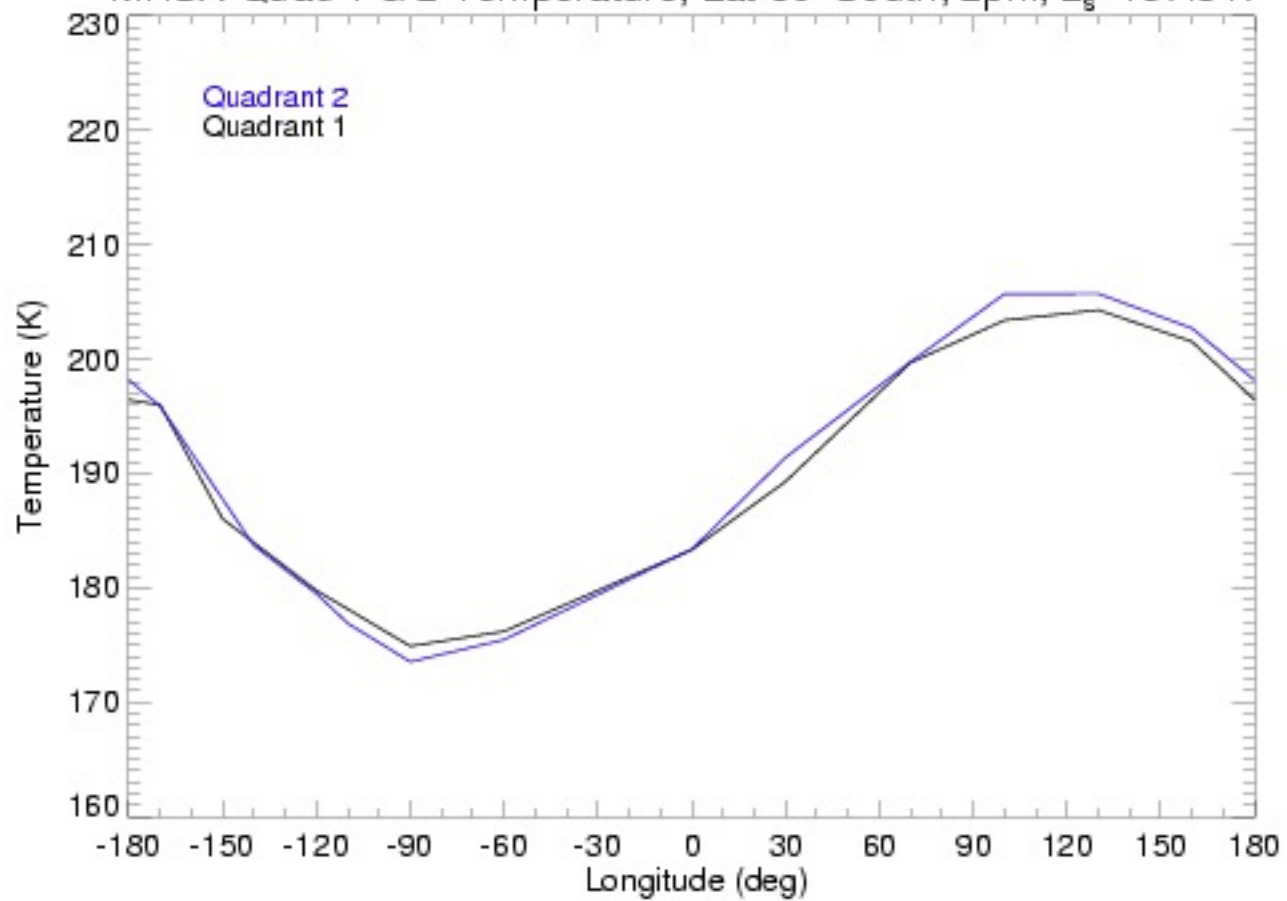
Regional storm on the southwestern rim of Hellas, $L_s=181^\circ$



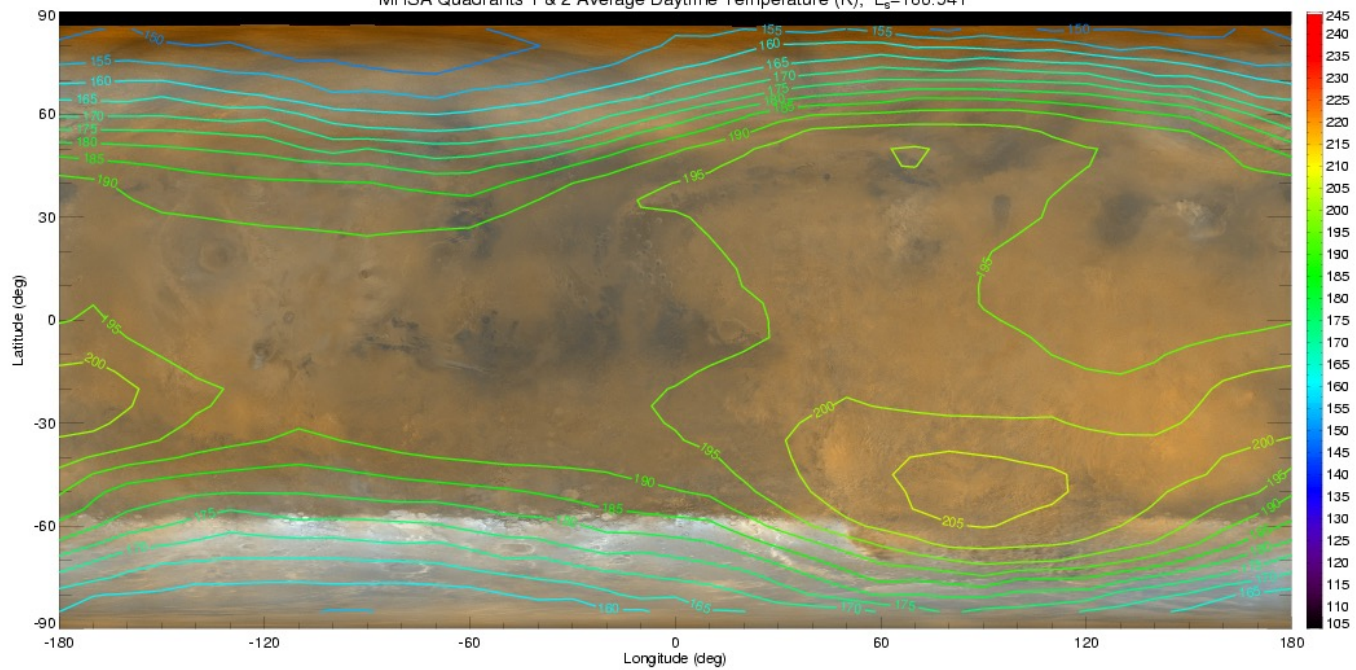
Development and propagation of wave one feature: $L_s=183^\circ-L_s=190^\circ$

- Maximum amplitude occurs at $\sim L_s=187^\circ$
- 30K - 40K peak to trough temperature change over latitude circle of 60°
- Hemispheric symmetry of both excitation and relaxing back to a more zonally uniform state (happens semi-simultaneously).
- Hemispheric symmetry of the wave one feature may indicate excitation of a global mode.
- Less than 1° of L_s after maximum amplitude, dust lifting begins in the Claritas region.

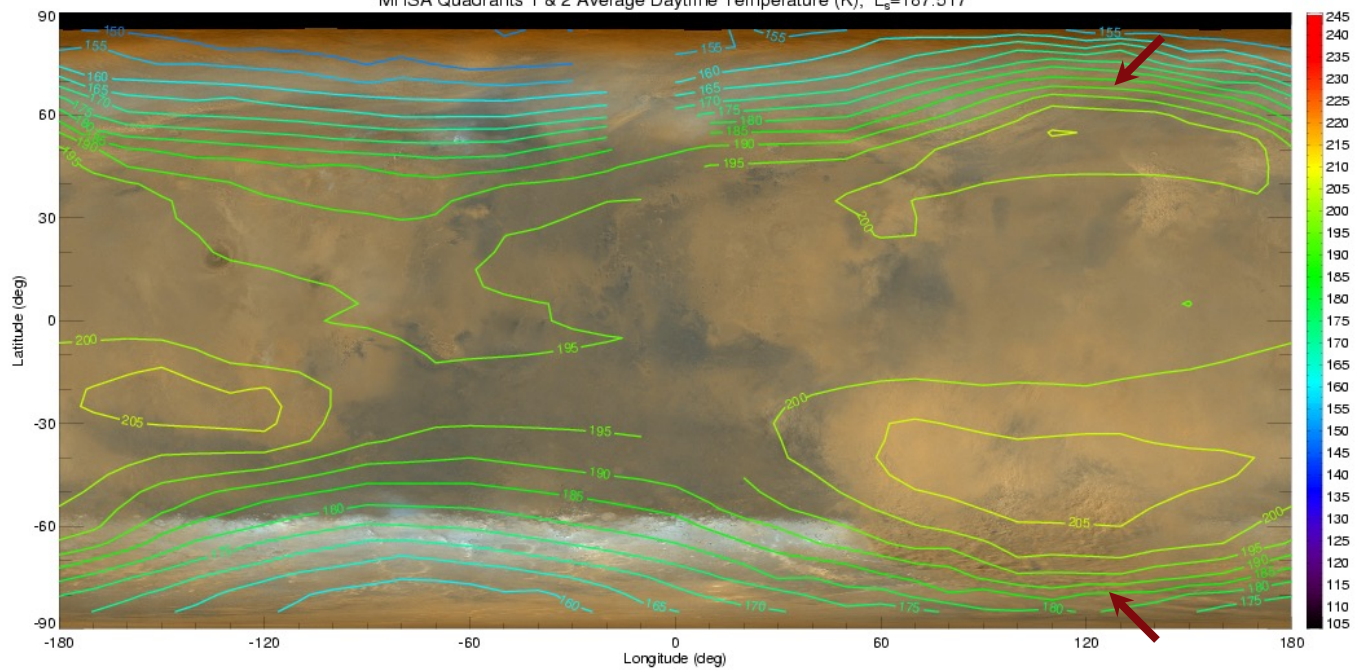
MHSA Quad 1 & 2 Temperature, Lat 60° South, 2pm, $L_s=187.517^\circ$



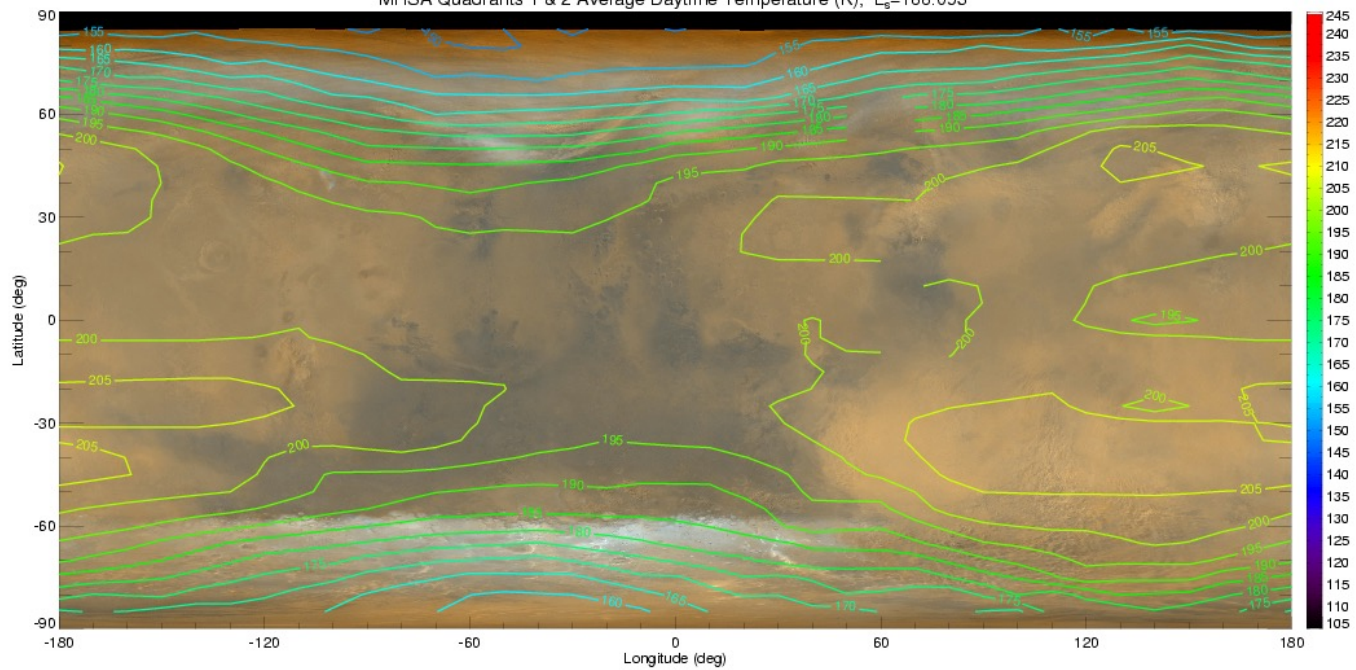
MHSA Quadrants 1 & 2 Average Daytime Temperature (K), $L_s=186.941^\circ$



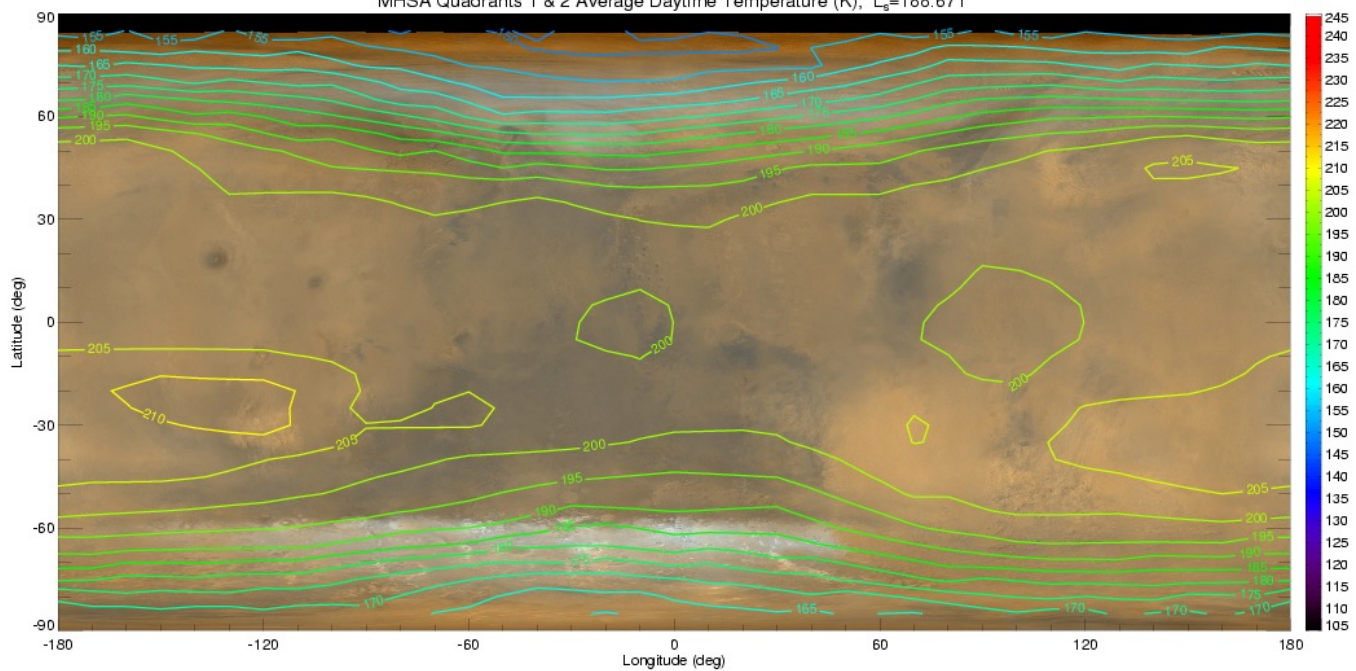
MHSA Quadrants 1 & 2 Average Daytime Temperature (K), $L_s=187.517^\circ$



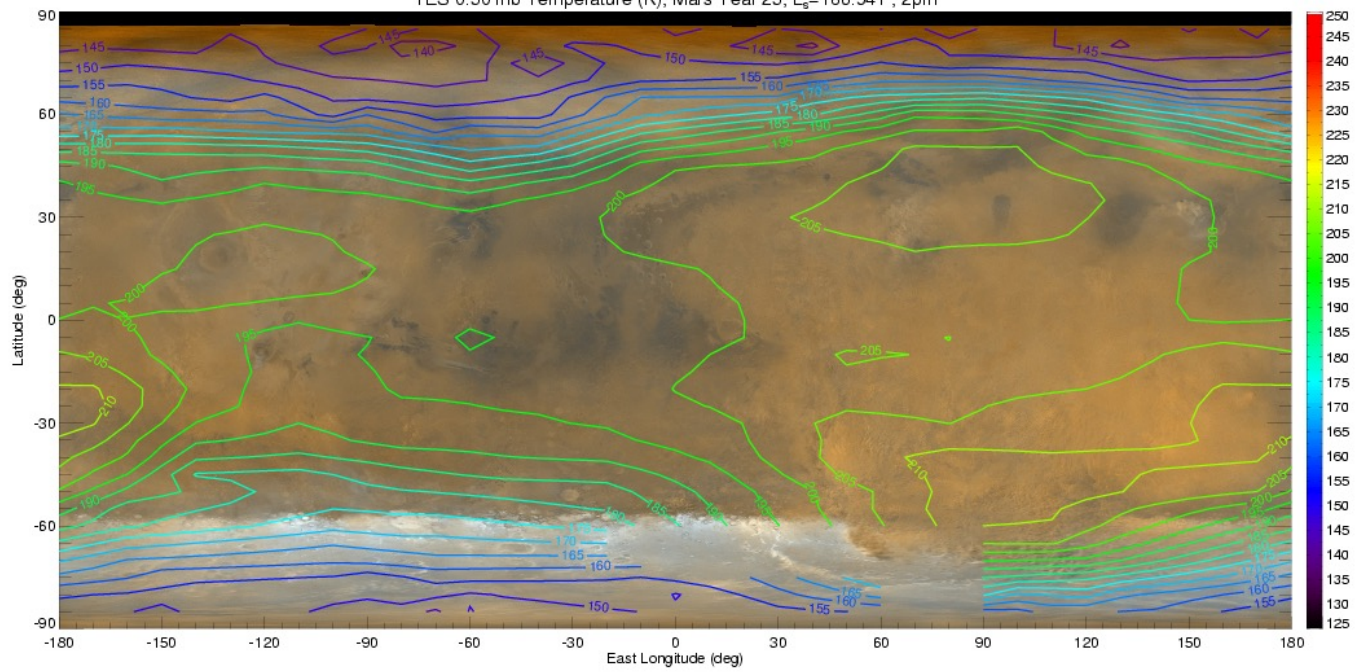
MHSA Quadrants 1 & 2 Average Daytime Temperature (K), $L_0=188.093^\circ$



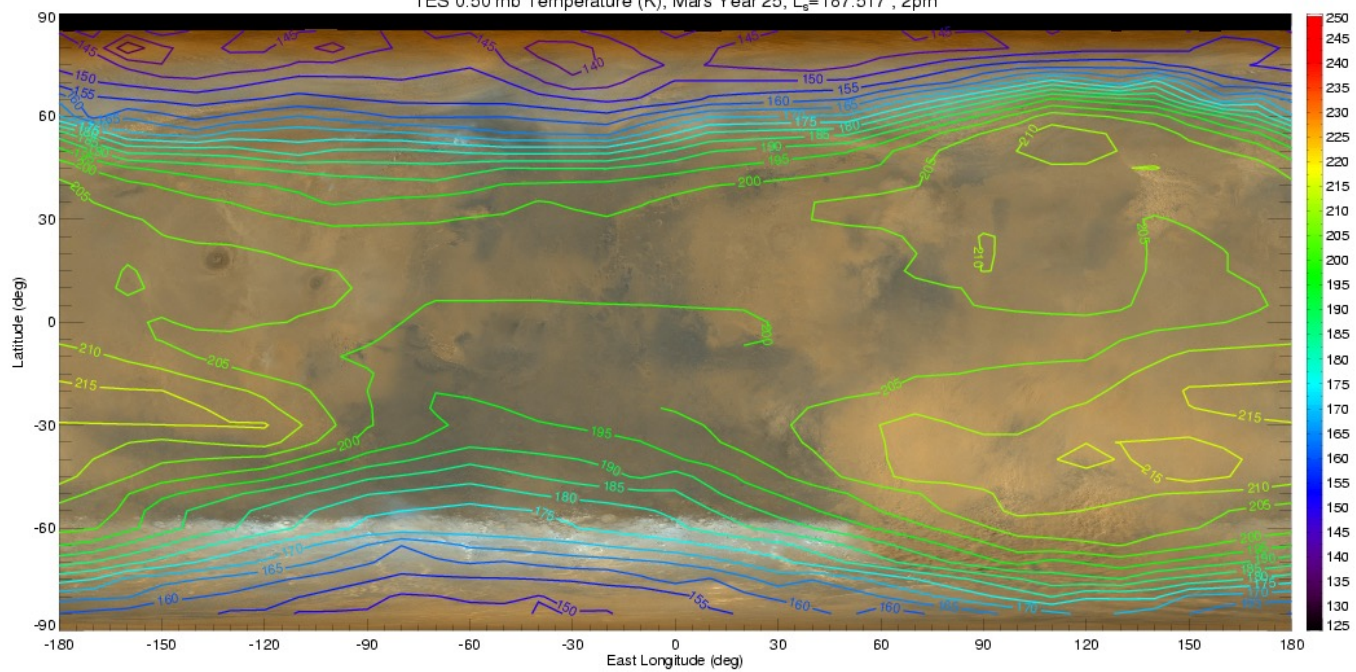
MHSA Quadrants 1 & 2 Average Daytime Temperature (K), $L_0=188.671^\circ$



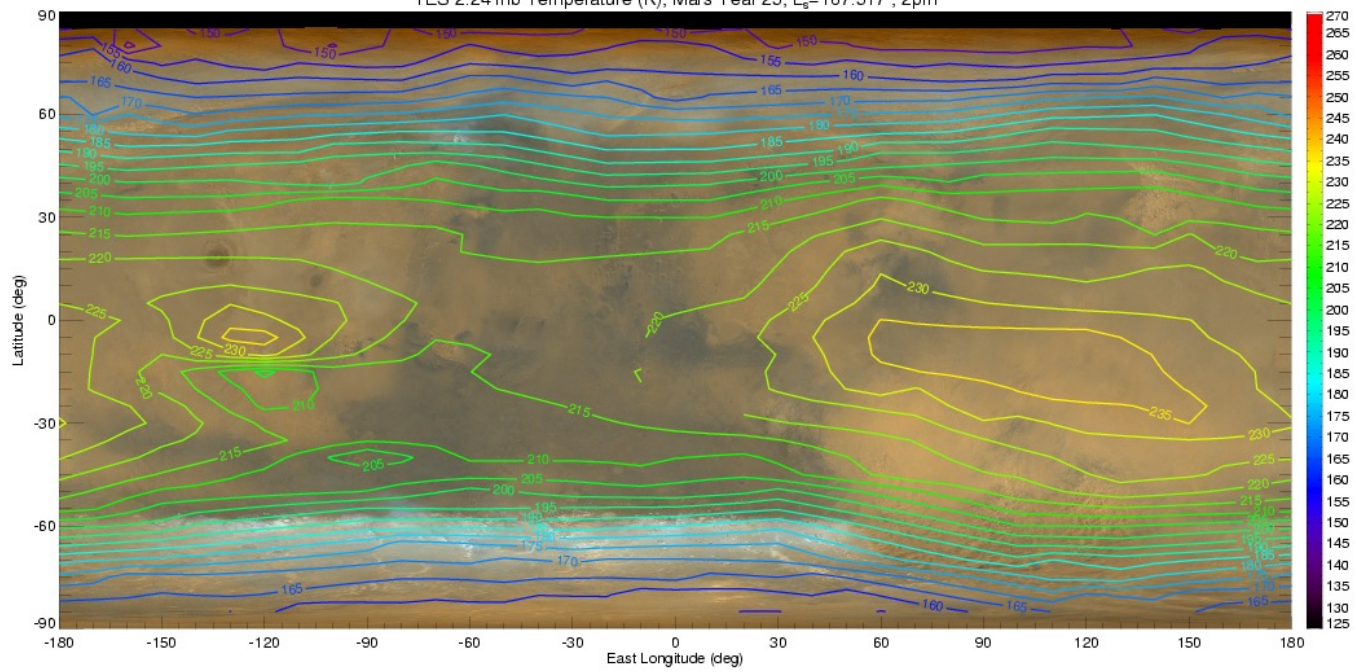
TES 0.50 mb Temperature (K), Mars Year 25, $L_s=186.941^\circ$, 2pm



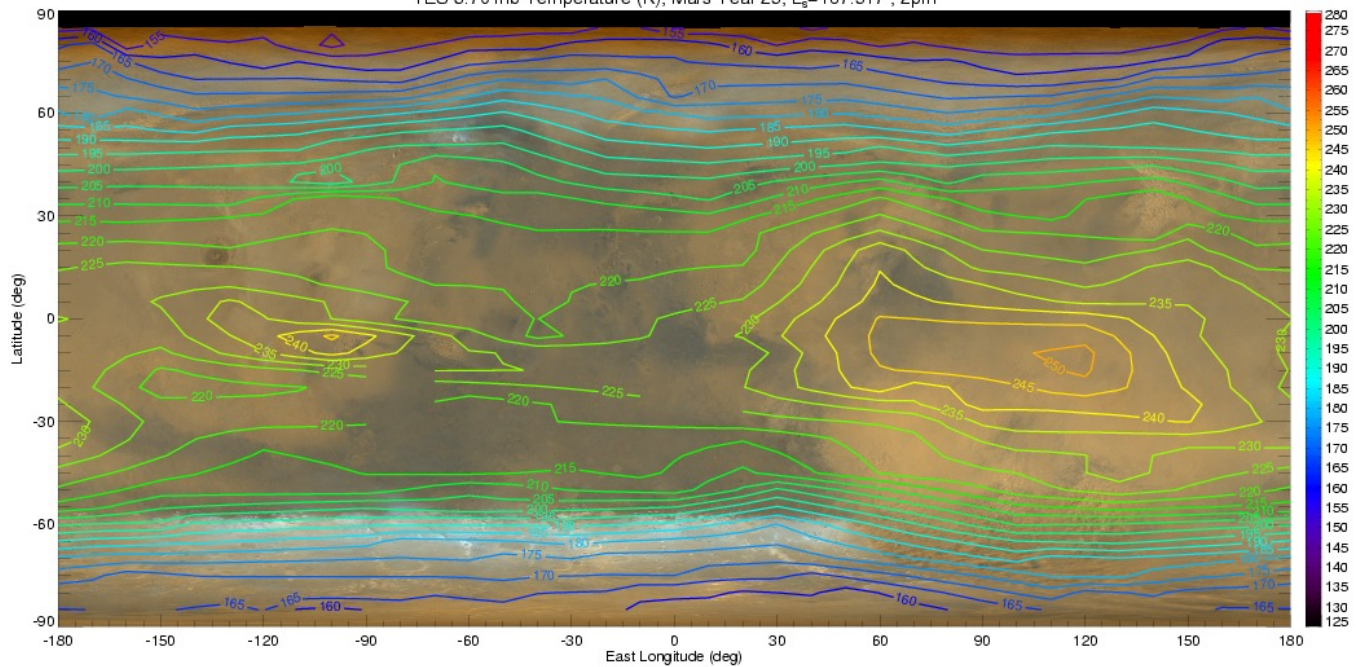
TES 0.50 mb Temperature (K), Mars Year 25, $L_s=187.517^\circ$, 2pm



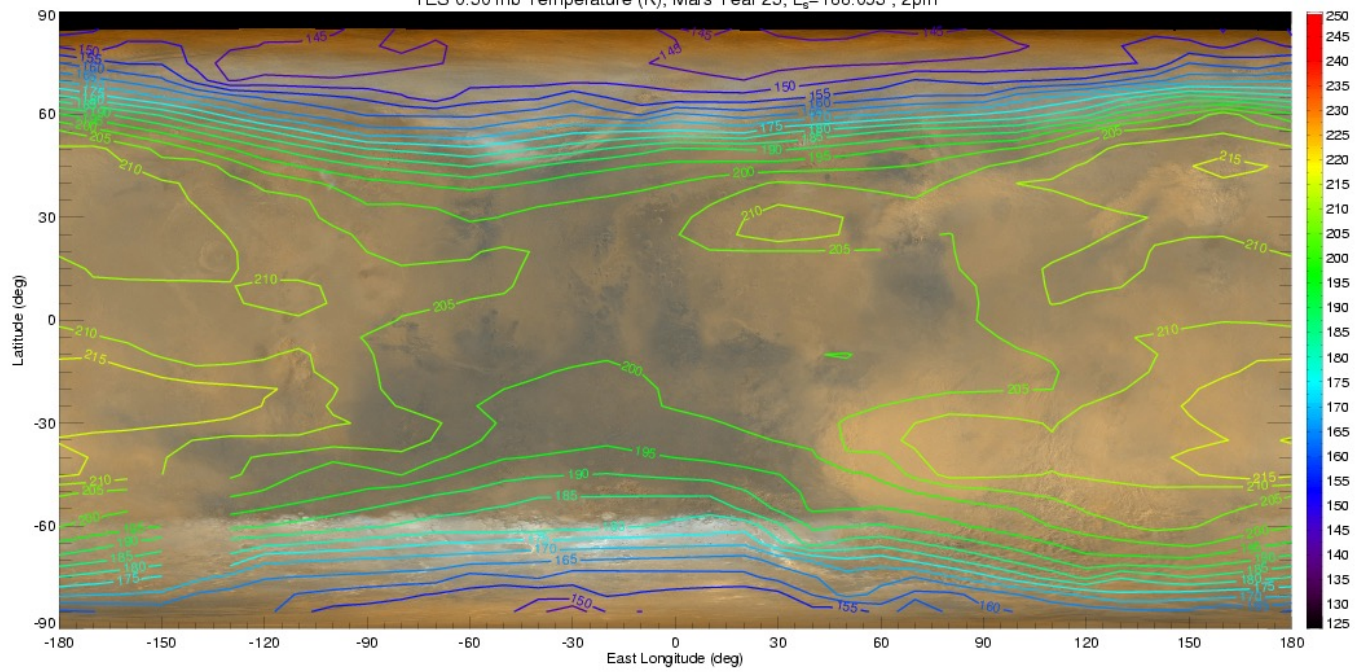
TES 2.24 mb Temperature (K), Mars Year 25, $L_s=187.517^\circ$, 2pm



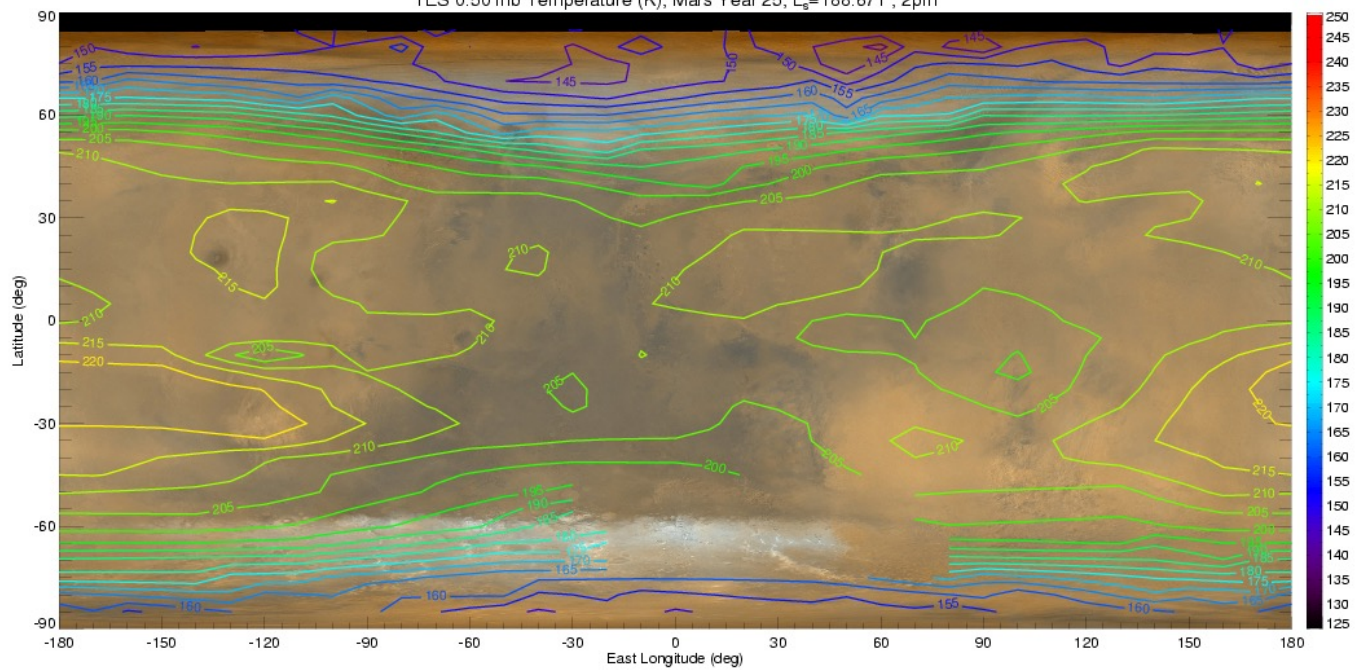
TES 3.70 mb Temperature (K), Mars Year 25, $L_s=187.517^\circ$, 2pm



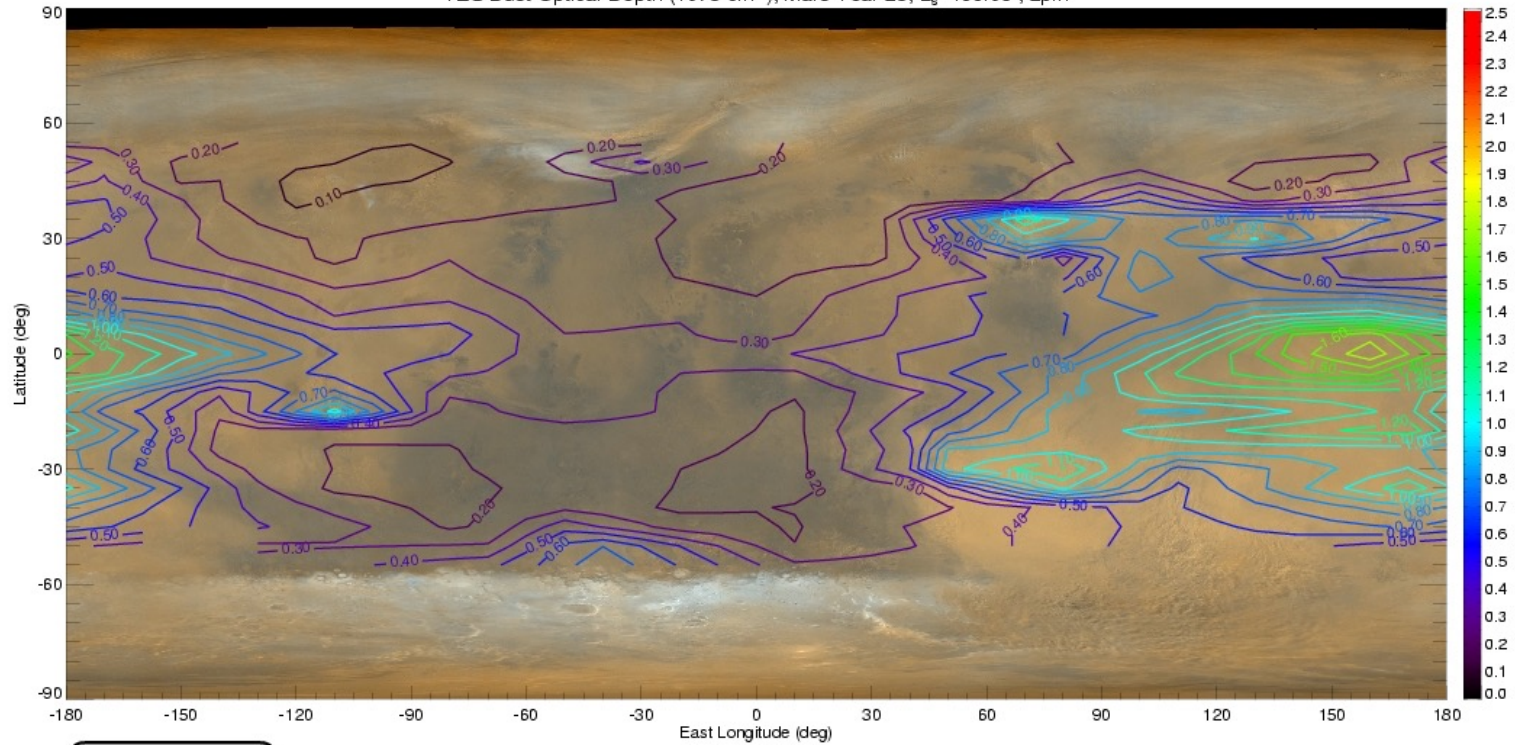
TES 0.50 mb Temperature (K), Mars Year 25, $L_0=188.093^\circ$, 2pm



TES 0.50 mb Temperature (K), Mars Year 25, $L_0=188.671^\circ$, 2pm



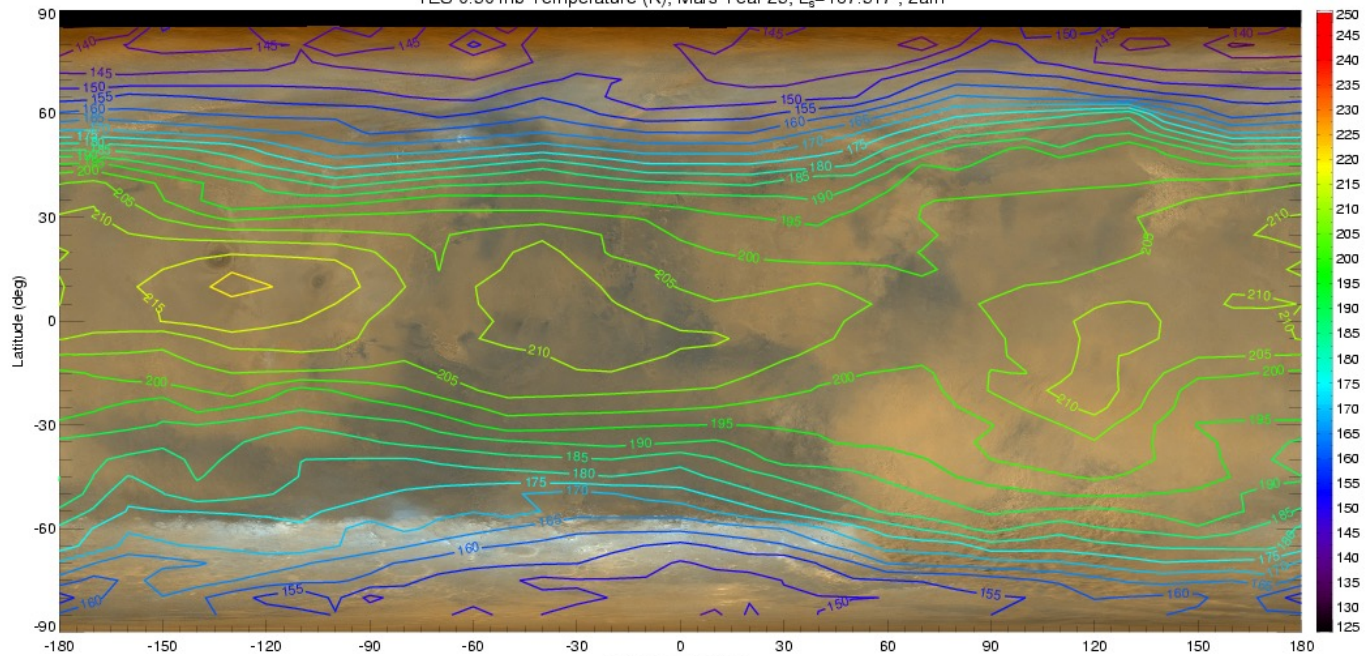
TES Dust Optical Depth (1075 cm^{-1}), Mars Year 25, $L_s=188.09^\circ$, 2pm



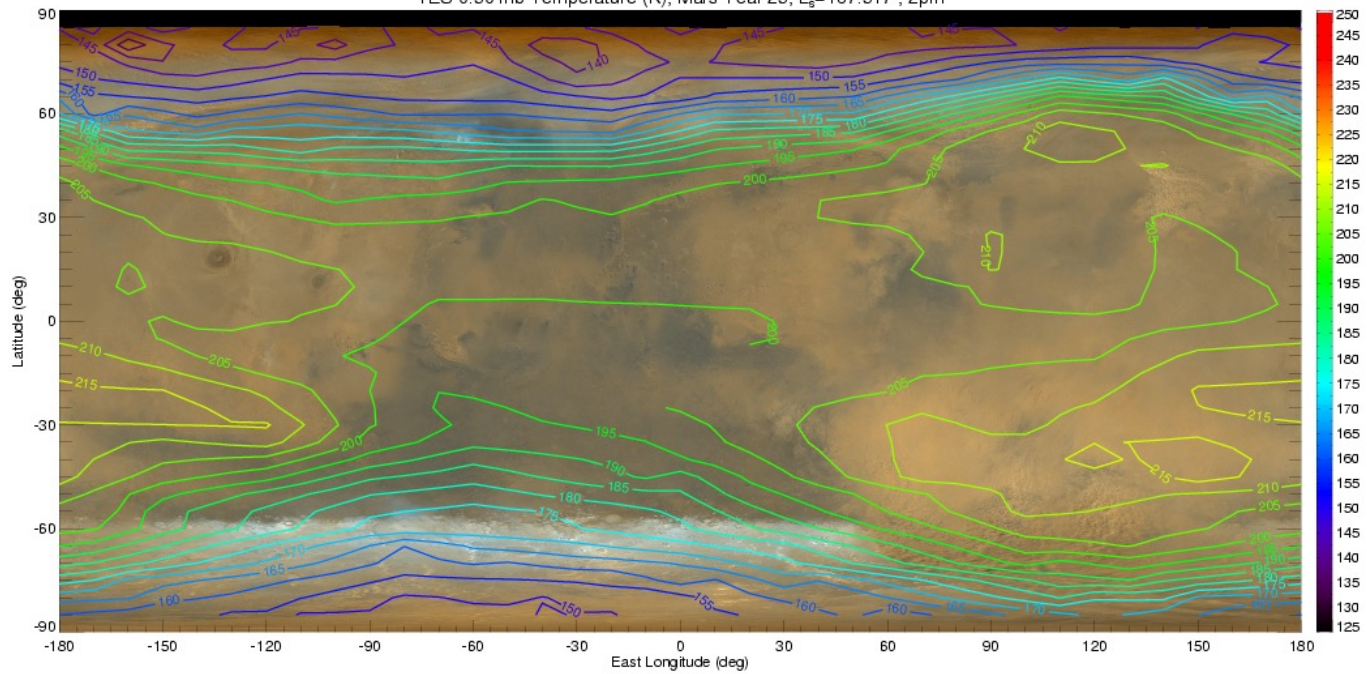
Dynamical processes

- Hemispheric symmetry is quite pronounced. What dynamical processes caused this wave one amplification?
- If it were purely a function of a diurnal tide, then one would expect both peak and trough to shift 180° from 2am to 2pm. Such a shift, however, is not evident in the observations.

TES 0.50 mb Temperature (K), Mars Year 25, $L_s=187.517^\circ$, 2am



TES 0.50 mb Temperature (K), Mars Year 25, $L_s=187.517^\circ$, 2pm



Working Hypothesis

- Precursory storm pulses are associated with traveling baroclinic eddies
- A combination of baroclinic eddies and cap edge flow triggered lifting in Hellas
- Once significant amounts of dust start to fill the atmosphere, wave amplification starts to occur.
- Southward movement of the dust is associated with a stationary wave
- Regional atmospheric heating in Hellas excites global Rossby wave train
- Claritas lifting is associated with propagation of Rossby wave train

Future Work

- Examine inter-annual variability: Look for wave features in subsequent years.
- Use the model:
 - First, prescribe the observed distribution of dust, adjust the altitude, and see what it excites.
 - Next, specify a source of dust in Hellas, and see what a truly consistent time interactive simulation would produce.
- Compare other lifting centers to Rossby wave train centers from model.